

Nuclear Chemistry

NOT Nuclear chemistry!

Types of radioactive decay

- Alpha (α) decay – the nucleus emits a helium nucleus (α particle).
 - $^{226}_{88}\text{Ra} \rightarrow ^4_2\text{He} + ^{222}_{86}\text{Rn}$
 - low energy process
 - can be blocked by skin or clothing or paper
 - can be dangerous if inhaled or ingested

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Types of radioactive decay

- Beta (β) decay – the nucleus emits an electron (β particle)
 - $^{14}_6\text{C} \rightarrow ^{14}_7\text{N} + ^0_{-1}\text{e}$
 - essentially the same as a neutron converting into a proton
 - $^1_0\text{n} \rightarrow ^1_1\text{p} + ^0_{-1}\text{e}$
 - higher in energy than α decay
 - can be blocked by heavy clothing or plastic
 - can also refer to the emission of a positron (positively charged electron or anti-electron)

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Types of radioactive decay

- Gamma (γ) emission – an excited nucleus releases energy and goes to a lower energy state.
 - highest energy
 - need lead or concrete to block it

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Nuclear Equations

- Represent what happens in a nuclear reaction
- nucleon number and charge must be conserved
 - mass numbers on top must be equal on both sides
 - atomic numbers on bottom must be equal on both sides

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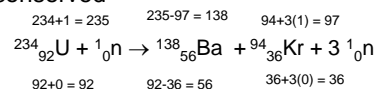
Nuclear Equations

- Reactant and product nuclei are represented by their nuclide symbols
 - $^{12}_6\text{C}$ or $^{241}_{94}\text{Pu}$
- Other symbols that may be used
 - proton – ^1_1H or ^1_1p or p
 - neutron – ^1_0n or n
 - electron – $^0_{-1}\text{e}$ or e or $^0_{-1}\beta$ or β
 - positron – ^0_1e or $^0_1\beta$
 - Gamma photon – $^0_0\gamma$

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“Balancing” nuclear equations

- There will be only one thing missing from the equation.
- Fill in the blank by making sure the nucleon number and charge are conserved



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Radiation Measurement

- Usually measured with a Geiger counter
- Unit is the Curie (Ci)
 - $1 \text{ Ci} = 3.7 \times 10^{10} \text{ disintegrations s}^{-1}$
 - being replaced by the Becquerel (SI Unit, Bq,
 $1 \text{ Bq} = 1 \text{ dis/s}$)

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Biological Effect

- rad (radiation absorbed dose) – amount of radiation absorbed by 1 g of tissue.
 - being replaced by the Gray (SI Unit, Gy, $1 \text{ Gy} = 100 \text{ rad}$)
- rem (radiation equivalent in humans)
 - being replaced by the Sievert (SI Unit, Sv, $1 \text{ Sv} = 100 \text{ rem}$)
- Actual dose (rem) = rad \times factor

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Factors

- For β and γ radiation the factor is 1
 - rads and rems are the same
- For α radiation the factor is 20.
- For high speed protons and neutrons the factor is 10.

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Radiation Exposure

- Always exposed to radiation
 - natural sources – rocks, cosmic sources
- “Artificial” sources
 - X-rays, CAT scans, PET scans
 - TV
 - Cigarettes

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Radiation Sickness

- Too much radiation
- Whole body exposure
 - 100 rem – temporary decrease in white blood cells
 - >100 rem – nausea, vomiting, fatigue, white blood cell reduction
 - 300 rem – white blood cell count is zero, diarrhea, hair loss, and infection
 - 500 rem – death in 50% of people exposed
 - 600 rem – all humans exposed are dead within a few weeks

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half-life of radioactive isotopes

- $\frac{1}{2}$ life is the amount of time it takes for $\frac{1}{2}$ of the nuclei present to decay
 - This number is a constant – it doesn't depend on the amount of substance
 - Example – ^{51}Cr has a half-life of 28 days
 - 100 g at 0 days
 - 50 g at 28 days
 - 25 g at 56 days
 - 12.5 g at 84 days

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Dating objects by half-lives

- ^{14}C has a half-life of 5730 years.
- Can be used to date object that take in carbon – i.e., living things
- Not useful past about 50,000 years

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Medical Applications of Radioactivity

- PET scans – Positron Emission Tomography
 - Uses a positron emitter (^{11}C , ^{15}O , and ^{13}N)
 - ingested by the patient and then the machine "sees" where the positrons come from and makes a picture.

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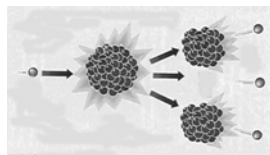
Fusion and Fission

- Nuclear fusion
 - A nucleus breaks apart and releases energy
 - Application of Einstein's equation $E=mc^2$
 - Process used in nuclear reactors
 - controlled reaction
 - Also used in nuclear weapons (atomic bombs)
 - chain reaction

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Chain Reaction

- One neutron causes a nucleus to split and it produces 2 or more neutrons
- Each of those does the same thing
- and so on
- and so on
- and so on



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Fusion and Fission

- Nuclear Fusion
 - Lighter nuclei are forced together to make heavier nuclei
 - Process that powers the sun and all the other stars
 - Requires high temperatures (100,000,000K) and pressures

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